identically, and relation (4.2) takes the form

$$
\begin{equation*}
I_{1}\left(r_{2}-r_{3}\right)+I_{2}\left(r_{3}-r_{2}\right)+I_{2}\left(r_{1}-r_{2}\right)=0 \tag{4.5}
\end{equation*}
$$

The above relation certainly holds in the first, as well as in the second case of integrability, though the conditions for a complementary integral to exist in these cases $\left(r_{1}=\right.$ $r_{2}=r_{3}$ or $I_{1}=I_{2}, r_{1}=r_{2}$ ) are not necessary for the relation to hold. We also note that relation (4.5) does not hold for a homogeneous ellipsoidal body.
5. In the case of an arbitrarily small perturbation in the integrable problem of the motion of a sphere whose centre of mass coincides with its geometrical centre, the necessary conditions for a complementary first integral to exist, analytic in the phase variables, in the class of bodies with ellipsoidal surfaces ( $a_{k}$ and $r_{k}$ in (2.1) are such, that $a_{1}{ }^{2}+a_{2}{ }^{2}+$ $\left.a_{3}{ }^{2} \neq 0, r_{1}{ }^{2}+r_{2}{ }^{2}+r_{3}{ }^{2} \neq 0\right)$, are combinations of the corresponding conditions of Sect. 3 and 4. This proves the following

Theorem. The following three conditions are simultaneously necessary for a complementary first integral to exist, analytic in the phase variables, in the problem of the motion of a heavy rigid ellipsoidal, nearly spherical body, whose centre of mass coincides with its geometrical centre and the moments of inertia are all different: l) the centre of mass of the ellipsoid coincides with its geometrical centre; 2) the principal axes of the inertia ellipsoid and surface ellipsoid coincide; 3) the moments of inertia of the ellipsoid and the semi-axes of its surface are connected by the relation

$$
I_{1}\left(\rho_{2}-\rho_{3}\right)+I_{2}\left(\rho_{3}-\rho_{1}\right)+I_{3}\left(\rho_{1}-\rho_{2}\right)=0
$$

The problem of the existence of a complementary analytic integral in the problem of the motion of a body of arbitrary, nearly spherical shape, whose centre of mass coincides with its geometrical centre, is more interesting and more complex. In this case, the first approximation in terms of a small parameter already yields a potential which may represent, generaliy speaking, an arbitrary function of the direction cosines $\gamma_{1}, \gamma_{3}, \gamma_{0}$, unlike the function $H_{1}$ (2.2) representing the sum of the linear and quadratic forms of the variables $\gamma_{1}, \gamma_{2}, \gamma_{3}$.

## REFERENCES

1. APPEL' P. Theoretical Mechanics. Vol.2, Moscow, Fizmatgiz, 1960.
2. KARAPETYAN A.V., On the stability of the steady motions of a heavy rigid body on a perfectly smooth horizontal plane. PMM Vol.45, No. 3, 1981.
3. ZIGIIN S.L., Splitting of the separatrices, branching of the solutions anc non-existence of an integral in the dynamies of a rigid body. Tr. Mosk. matem. o-va, vol.41, 1980.
4. KOZLOV V.V. and ONISHCHENKO D.A., Non-integrability of Kirchhoff's equations. Dokl. Akad. Nauk SSSR, Vol. 26E, NC.6, 1982.

Translated by L.K.

PMM U.S.S.R.,VOL.49,No. 3,pp. 389-392,1985
0021-8928/85 \$10.00+0.00
Printed in Great Britain
Pergamor. Journals Itd.

# PERTURBED MOTION OF A KOVALEVSKAYA TOP* 

## N.N. MOTORINA

Perturbation theory based on the appiication of Lie series, is used to study a special case of the motion of a rigid body about a fixed point. The equations written in action-angle variables are used in Hamiltonian form. The solutions are obtained in the form of trigonometric series with constant coefficients.
It is assumed that the distribution of mass in the body is clsoe to the distribution in the Kovalevskaya case and the centre of gravity of the body is situated fairly near to the fixed point. The canonical Deprit variables/ / / are used. The motion of the body can be described in these variables by the following set of equations:

$$
\frac{d(L, G, H)}{d t}=\frac{\partial F}{\partial(l, \xi, h)}, \quad \frac{d(l, g, h)}{d t}=-\frac{\partial F}{\tilde{d}(L, G, H)}
$$

Using the condition that the centre of gravity of the body is situated fairly close to the fixed point and the principal moments of inertia $A$ and $B$ differ from each other, we can
write the Hamiltonian of these equations, in accordance with the order of magnitudies, in the form /2/

$$
\begin{aligned}
F & =F_{0}+F_{1} \\
F_{0} & =\frac{G^{2}-L^{2}}{2 A}-\frac{L^{2}}{2 C} \\
F_{1} & =\frac{A-B}{2 A B}\left(G^{2}-L^{2}\right) \cos ^{2} l+\frac{p x_{c}}{G^{2}}\left(L \sqrt{G^{2}-H^{2}} \sin l \cos g+\right. \\
& \left.G \sqrt{G^{2}-H^{2}} \cos l \sin g+H \sqrt{G^{2}-L^{2}} \sin l\right)
\end{aligned}
$$

Here $A, B, C$ are the principal moments of inertia for the fixed point, $P$ is the weight of the body and $x$ is the coordinate of the centre of mass in the principal axes of inertia.

We take $\varepsilon=\max \left\{A-B, x_{c}\right\}$ as the small parameter. To solve the equationswith the Hamiltonian (1) it is convenient to apply the method of the theory of canonical transformations using Lie series, developed by Hori /3/.

We eliminate the variable $l$ from the Hamiltonian by carrying out the canonical variable transformation

$$
L, G, H, t, g, h \rightarrow L^{\prime}, G^{\prime}, H^{\prime}, l^{\prime}, g^{\prime}, h^{\prime}
$$

With help of the generating function $S^{\prime}\left(L^{\prime}, G^{\prime}, H^{\prime}, l^{\prime}, g^{\prime}, h^{\prime}\right)=S_{2}^{\prime}+S_{i}^{\prime}+\ldots$. We assume that the Hamiltonian of the new equations can also be written according to the order of magnitudes in the form $F^{\prime}=F_{0}{ }^{\prime}+F_{1}{ }^{\prime}+\ldots$.

According to the Hori method the components of the generating function and the Hamiltonian are obtained from the formulas

$$
\begin{align*}
& F_{0}^{\prime}=F_{0}, \quad F_{1}^{\prime}=F_{1 s t} \quad S_{1}^{\prime}=\int F_{12} d t^{\prime}, \quad F_{2}^{\prime}=F_{2 s}+1 / 2\left\{F_{1}+F_{1}^{\prime}, S_{1}^{\prime}\right\}_{s *}  \tag{2}\\
& S_{2}^{\prime}=\int\left|F_{2 F}-1_{2}^{\prime}\left\{F_{1} \div F_{1}^{\prime}, S_{2}^{\prime}\right\}_{p}\right| d t^{\prime}
\end{align*}
$$

etc. The curly brackets are the poisson brackets and the indices $p$ and s denote the periodic and secular part in $t$ respectively.
the parameter $t$ ' is introduced by means of the equations

$$
\frac{d\left(L^{\prime} \cdot G^{\prime} \cdot H^{\prime}\right)}{d t}=\frac{\partial F_{0}^{\prime}}{\partial\left(l, g^{\prime}, h\right)}, \quad \frac{d\left(l^{\prime} \cdot \mathrm{E}^{\prime}, h^{\prime}\right)}{d t^{\prime}}=-\frac{\partial F_{0}^{\prime}}{\partial\left(L^{\prime}, G^{\prime}, H^{\prime}\right)}
$$

The following poisson bracket holds by virtue of these equations:

$$
\left\{F_{0}, S_{k}\right\}=-d s_{k} / d t^{\prime}
$$

and this gives the relation connectingt'and $l$

$$
\begin{equation*}
d t^{\prime}=-\left(\partial F_{0} \cdot \partial L\right)^{-1} d t^{\prime} \tag{3}
\end{equation*}
$$

Applying algorithm (2) to the initial Hamiltonian $F$, taking (3) into account, we obtain (the prime accompanying the variables are omitted for convenience)

$$
\begin{align*}
& F_{0}{ }^{\prime}=\frac{G^{2}-L^{2}}{2 A}-\frac{L^{2}}{2 C}, \quad F_{1}=\frac{A-B}{44 B}\left(G^{2}-L^{2}\right)  \tag{4}\\
& s_{1}^{\prime}=-\frac{C(A-B)}{8 E L(A-C)}\left(C^{2}-L^{2}\right) \sin 2 l-\frac{P A C X}{L G^{2}(A-C)} \\
& \left(L \sqrt{G^{2}-H^{2}} \cos l \cos g-G \sqrt{G^{2}-H^{2}} \sin l \sin g+H \sqrt{G^{2}-L^{2}} \cos l\right) \\
& F_{z^{\prime}}=\left[\frac{C(A-B)^{2}\left(G^{2}-L^{2}\right)\left(G^{2}+3 L^{2}\right)}{64 A B^{2} L^{2}(A-C)}+\frac{k H^{2}}{4 G^{2}} \sqrt{\frac{G^{2}-H^{2}}{G^{2}-L^{2}}}-\frac{k H^{2}}{2 G^{2}}+\right. \\
& \left.\frac{k}{4 L^{2}}\right]-\frac{k\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right)}{4 L^{2} G^{4}} \cos ^{2} g-\frac{3 k H}{4 L G^{2}} \sqrt{\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right)} \cos g- \\
& \frac{k H\left(G^{2}-H^{2}\right)}{4 L G^{4}} \cos g, \quad k=\frac{P^{2} A C X_{r}{ }^{2}}{A-C}
\end{align*}
$$

The resulting hamiltoniar $F^{\prime}$ does not contain the variable $l^{\prime}$. To eliminate the remaining angle variable $g^{\prime}$ we carry out another canonical transformation

$$
L^{\prime}, G^{\prime}, H^{\prime}, l^{\prime}, g^{\prime}, h^{\prime} \rightarrow L^{*}, G^{*}, H^{n}, l^{*}, g^{*}, h^{*}
$$

with help of the generating function $S^{\prime}\left(L^{n}, G^{\prime \prime}, H^{n}, l^{n}, g^{n}, h^{\prime \prime}\right)=S_{1}{ }^{*}+S_{2}^{*}+\ldots$. As before, we assume that the Hamiltonian of the new equations of motion has the form $F^{\prime \prime}=F_{0}^{\prime \prime}+F_{1}^{*}+\ldots$.

To find the components of this Hamiltonian and the generating function, we use the formulas

$$
\begin{equation*}
F_{0}^{*}=F_{0^{\prime}}^{\prime}, \quad F_{1}^{*}=F_{1}^{\prime}, \quad F_{2}^{*}=F_{2 s^{\prime}}, \quad S_{1}^{*}=\int F_{2 p^{\prime}} d t^{n} \tag{5}
\end{equation*}
$$

etc. We introduce the parameter $t$ " by means of the equations

$$
\frac{d\left(L^{*}, G^{n}, H^{*}\right)}{d t^{*}}=\frac{\partial F_{1}^{*}}{\partial\left(l^{n}, g^{*}, h^{*}\right)}, \quad \frac{d\left(I^{n}, g^{*}, h^{n}\right)}{d t^{*}}=-\frac{\partial F_{1}^{*}}{\partial\left(L^{*}, G^{*}, H^{n}\right)}
$$

Then the relation connecting $t^{\prime \prime}$ with $g^{\prime \prime}$ will be
$d t^{*}=-\left(\partial F_{1}{ }^{\prime} / \partial G^{\prime \prime}\right)^{-1} d g^{n}$
Changing in expressions (5) from integration with respect to $t^{\prime \prime}$ to integration with respect to $g^{\prime \prime}$, we obtain (omitting the double primes accompanting the variables)

$$
\begin{aligned}
& F_{2^{\prime \prime}}=[\ldots]-\frac{k\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right)}{8 L^{2} G^{6}} \\
& S_{1^{\prime \prime}}=\frac{k_{1}\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right)}{8 L^{2} G^{6}} \sin 2 g+\frac{k_{1} H\left(G^{2}-H^{2}\right)}{2 L G^{8}} \sin g- \\
& \quad \frac{3 k_{1} H}{2 L G^{2}} \sqrt{\left.\frac{\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right.}{}\right)} ; \quad k_{1}=\frac{k A B}{A-C}
\end{aligned}
$$

The expressions for $F_{0}{ }^{*}, F_{1}{ }^{\prime}$ are identical with those given in (4), and the symbol [...] denotes the expression within the square brackets for $F_{i}$ in (4).

The Hamiltonian $F^{\prime \prime}$ obtained contains no angle variables. The equations of motion yield directly $L^{\prime \prime}=$ const, $G^{\prime \prime}=$ const, $H^{\prime \prime}=$ const, while the angle variables will be linear functions of time

$$
l^{*}=l^{\wedge} t+l_{0}, \quad g^{n}=g^{\prime \prime} t+g_{0}, \quad h^{n}=h^{n} t+h_{0}
$$

where $l_{0}, g_{0}, h_{0}$ are the initial values of the corresponding varaibles. From the equations of motion we obtain (double primes accompanying the variables are omitted)

$$
\begin{aligned}
& r^{*}=-\frac{L(A-C)}{A C}+\frac{L(A-B)}{2 A B}+\frac{C(A-B)^{2}\left(G^{4}+3 L^{4}\right)}{32 A B^{2} L^{3}(A-C)}- \\
& \frac{k\left(G^{2}+H^{2}\right)}{4 G^{2} L^{3}}-\frac{k L H^{2}}{4 G^{4}\left(G^{2}-L^{2}\right)} \sqrt{\frac{G^{2}-H^{2}}{G^{2}-L^{2}}}+O\left(\varepsilon^{3}\right) \\
& g^{\prime \prime}=-\frac{G(A+B)}{2 A B}-\frac{C G(A-B)^{2}\left(G^{2}+L^{2}\right)}{16 A B^{2} L^{2}(A-C)}-\frac{k\left(G^{2}-10 H^{2}\right)}{4 G^{5}}-\frac{k H^{2}}{4 L^{2} G^{3}}+ \\
& \frac{k H^{2}\left[G^{2}\left(4 G^{2}-3 L^{2}\right)-H^{2}\left(5 G^{2}-4 L^{2}\right)\right]}{G^{5}\left(G^{2}-L^{2}\right) \sqrt{\left(G^{2}-H^{2}\right)\left(\bar{G}^{2}-L^{2}\right)}}+O\left(\varepsilon^{3}\right) \\
& h^{n}=\frac{5 k H}{4 G^{4}}-\frac{k H}{4 L^{2} G^{2}}-\frac{k H\left(2 G^{2}-3 H^{2}\right)}{4 G^{4} \sqrt{\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right)}}+O\left(\varepsilon^{3}\right)
\end{aligned}
$$

The last expression contains only the terms of the second order of smallness in the small parameter.

The initial variables and any function of the initial variables can be found using the formula

$$
\begin{align*}
& f(L, G, H, l, g, h)=f\left(L^{\prime \prime}, G^{n}, H^{n}, l^{\prime \prime}, g^{\prime \prime}, h^{n}\right)+\left\{f, S^{\prime}+S^{*}\right\}+  \tag{6}\\
& \quad 1 / 2\left\{f,\left\{S^{\prime}, S^{\prime \prime}\right\}\right\}+1 / 2\left\{\left\{f, S^{\prime}-S^{n}\right\} S^{\prime}-S^{\prime \prime}\right\} \div O\left(\varepsilon^{3}\right)
\end{align*}
$$

Thus the theory in qeustion implies that $H=H^{*}=$ const. If we write the terms up to and including the first-order terms in the small parameter, the expressions for the variables $L, G$ and $h$ will be (the double primes accompanying the variables are omitted)

$$
\begin{aligned}
& L= L-\frac{C(A-B)\left(G^{2}-L^{2}\right)}{4 B L(A-C)} \cos 2 l-\times\left[L \sqrt{G^{2}-H^{2}} \sin l \cos g+\right. \\
&\left.G \sqrt{G^{2}-H^{2}} \cos l \sin g+H \sqrt{G^{2}-L^{2}} \sin l\right]+O\left(\varepsilon^{2}\right) \\
& G=G-\times\left[L V \overline{G^{2}-H^{2}} \cos l \sin g+G \sqrt{G^{2}-H^{2}} \sin l \cos g\right]+ \\
& \frac{k_{1}\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right)}{4 L^{2} G^{6}} \cos 2 g+\frac{\left(k_{1} H\left(G^{2}-H^{2}\right)\right.}{2 L G^{3}} \cos g- \\
& \frac{3 k_{1} H}{2 L G^{5}} \sqrt{\left(G^{2}-H^{2}\right)\left(G^{2}-L^{2}\right)} \cos g+O\left(\varepsilon^{2}\right) \\
& h= h_{0}+\times\left[\frac{G H}{\sqrt{G^{2}-H^{2}}} \sin l \operatorname{sing+\sqrt {C^{2}-L^{2}}\operatorname {cos}l-}\right. \\
&\left.\frac{L H}{\sqrt{G^{2}-H^{2}}} \cos l \cos g\right]-\frac{k_{1} H\left(G^{2}-L^{2}\right)}{4 L^{2} G^{3}} \sin 2 g+ \\
& \frac{k_{1} H\left(G^{2}-3 H^{2}\right)}{2 L G^{5}} \sin g-\frac{3 k_{1}\left(G^{2}-2 H^{2}\right)}{2 L G^{5}} \sqrt{\frac{G^{2}-L^{2}}{G^{2}-H^{2}}} \sin g+O\left(\varepsilon^{2}\right) \\
& \times \frac{P A C X_{c}}{L G^{2}(A-C)}
\end{aligned}
$$

Hence, we obtain the values of these variables in the form of trigonometric functions whose constant coefficients depend on $L^{\prime \prime}, G^{\prime \prime}, H^{\prime \prime}$. The expressions for the angle variables $l$ and $g$ contain, in addition to the analogous terms, a secular part. It should be noted that when the theory is constructed including the second-order terms the secular part also appears in the expression for $h$.

Since the Poisson brackets are invariant, formula (6) enables us to determine, fairly simply, in the form of series of the same type, e.g. the variables $p, q, 2, \gamma, \gamma^{\prime}, \gamma^{\prime \prime}$ of the EulerPoisson problem, provided that we use their expressions in terms of the Deprit variables/4/ and the values obtained for the generating functions $S^{\prime}$ and $S^{\prime \prime}$. The latter depend only on the new variables. This significantly facilitates the computations and makes them suitable for computer use.

## REFERENCES

1. DEPRIT A., Free rotation of a rigid body studied in the phase plane. Amer. J. Phys., Vol. 35, No. 5, 1967.
2. DEMIN V.G. and KISELEV F.I., On the periodic motions of a rigid body in a central Newtonian field. PMM Vol. 38, No. 2, 1974.
3. HORI G. , Theory of general perturbetion with unspecified canonical variables. Publ. Astron. Soc. Japan, Vol. 18, No. $4,1966$.
4. ARKHANGEL'SKII YU.A., Analytical Dynamics of a Rigid Body. Moscow, Nauka, 1977.

## ON THE STUDY OF RANDOM OSCILLATIONS IN NON-AUTONOMOUS MECHANICAL SYSTEMS USING THE FOKKER-PLANCK-KOLMOGOROV EQUATIONS*

## NGUEN DONG AN

A method of integrating the Fokker-Plank-Kolmogorov equations (Fpke: used
In the theory of random oscillations/l-4/is proposed. The Duffing
equetion is first studied as an example. The method is then used, together
with the method of averaging, to study random oscillations of non-autonomous
mechanical systems with one degree of freedom wher the eigenfreguency varies
in a random manner. The van-der-Pol equation is considered for the case
of a randomly varying eigenfrequency and periodic parametric excitation.
When the function sought is replaced, the FPKE transform into another
equation whose trivial solutions have the corresponding particular solutions
of the FPKE. The condition of integrability of the FPKE is obtained as
the direct consequence of the change in question.

1. Consider a mechanicai system with one degree of freedon, whose motion is described by the following stochastic equation:

$$
\begin{align*}
& x^{\prime \prime}-\omega^{2} z=f\left(x, z^{\prime}\right)-\sigma_{i}^{\prime}(t)  \tag{1.1}\\
& f\left(z, x^{\prime}\right)=\sum_{s=1}^{m} a_{;} \sum_{i, j=0}^{i-j=} \gamma_{i j} x^{i} x^{\prime}: \quad \alpha_{s^{\prime}} r_{i j}=\mathrm{const} \tag{1.2}
\end{align*}
$$

whexe for is a ranaom, winte noise-type action of unit intensity. using the sutstitution

$$
\begin{equation*}
x=a \cos \psi \cdot x^{*}=-a \omega \sin \psi \tag{1.2}
\end{equation*}
$$

and the Itc formila, we rejuce Eg. M. D to the form $/ 4 /$

$$
\begin{align*}
& d a=-\frac{1}{\omega} f(a \cos \psi,-a w \sin \psi) \sin \psi^{4}-\frac{5^{2}}{2 a \omega^{2}} \cos \psi^{2}, d t-  \tag{1,4}\\
& \frac{5}{\omega} \sin 4 d t(t) \\
& d \psi=\left[\omega-\frac{1}{a \omega} j(a \cos \psi .-a \omega \sin \psi) \cos \psi-\frac{\sigma^{2}}{a^{2} \omega^{2}} \sin \psi \cos \psi\right] a t- \\
& \frac{0}{a \omega} \cos \psi d(t)
\end{align*}
$$

Let us write the FPNE corresponaing to system (1.4) for the stationary probability density of the amplitude ano phase $W(a, \psi)$

$$
\begin{align*}
& \frac{\partial}{\partial a}\left[B_{1}(a, \psi) W\right]+\frac{\partial}{\partial \psi}\left[B_{2}(a, \psi) H\right]=\frac{1}{2}\left\{\frac{\partial^{2}}{\partial a^{2}}\left[B_{11}(a, \psi) H\right] \perp\right.  \tag{1,5}\\
& \left.2 \frac{\partial^{2}}{\partial a \partial \psi}\left[B_{12}(a, \psi) W\right]+\frac{\partial^{2}}{\partial \psi^{2}}\left[B_{22}(a, \psi) H\right]\right\}
\end{align*}
$$

Taking into account the expression for $f\left(x, x^{*}\right)(1,2)$, we obtain

$$
\begin{equation*}
B_{x}(a, \psi)=-\frac{\sin \psi}{\omega} f\left(a \cos \psi,-a(\omega \sin \varphi)+\frac{s^{z} \cos ^{z} \psi}{2 \omega^{2} a}=\frac{\theta^{2} \cos ^{2} \psi}{2 \alpha)^{2}} a^{-1}+\sum_{s=1}^{m} A_{s}(\psi) a^{*}\right. \tag{1,6}
\end{equation*}
$$

*PrikI.Matem.Mekhan., 49,3,506-512,1985

